

# β-delayed α-decay of <sup>16</sup>N

- experiments performed at the AGOR/Triµp-facility

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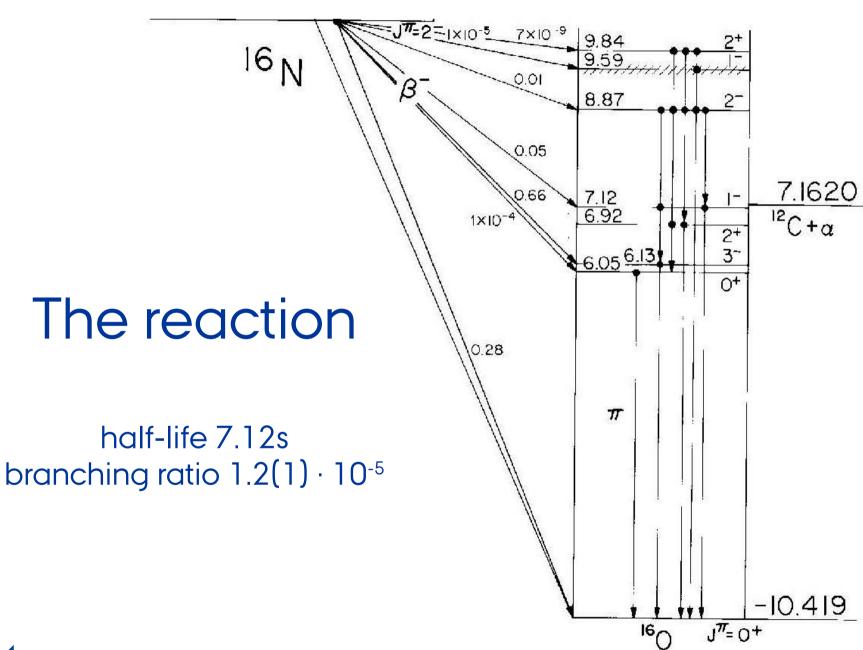
# **Physics**

•Explain relative abundance of <sup>12</sup>C and <sup>16</sup>O.

$$(3\alpha,\gamma)^{12}$$
C VS.  $^{12}$ C $(\alpha,\gamma)^{16}$ O

- Direct measurement of reaction at stellar He-burning energies is not feasible.
- E1-contribution to the reaction can be constrained by measuring the inverse process
- Large contribution to uncertainty comes from the branching ratio for  $\alpha$ -decay.

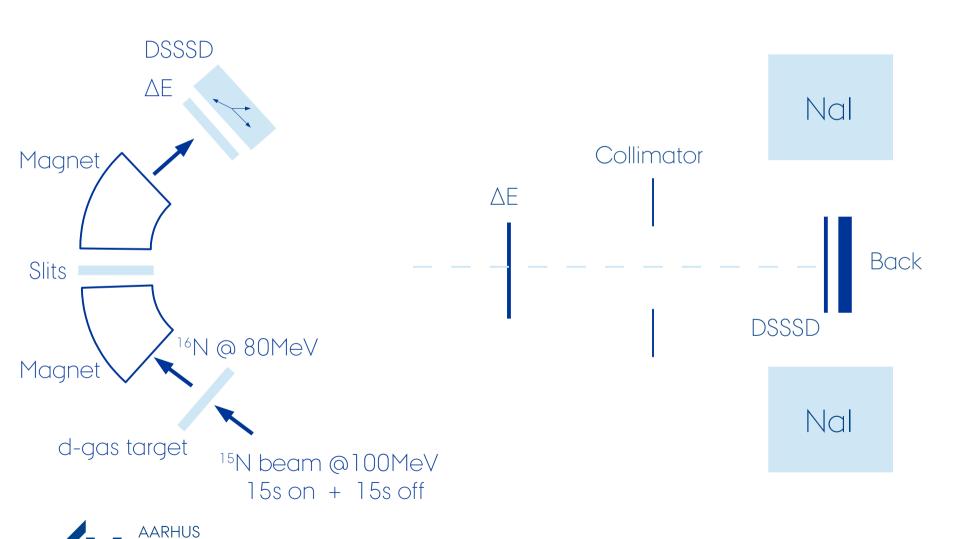






# Experiment



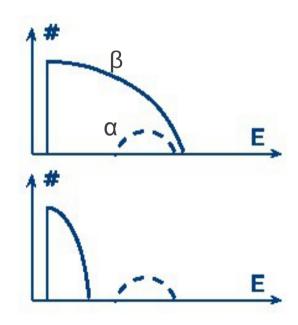


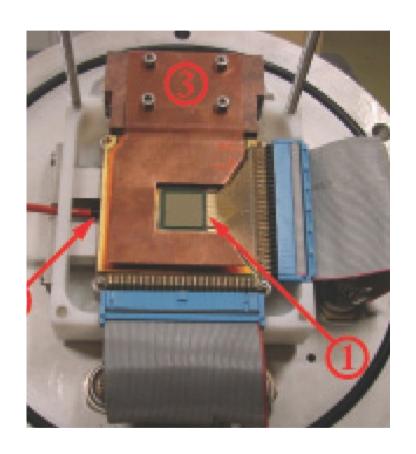
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#### Detector

- DSSSD, 48 x 48 strips, 2304 pixels.
- Pixels: 78 x 300 x 300 µm<sup>3</sup>.



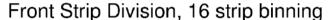


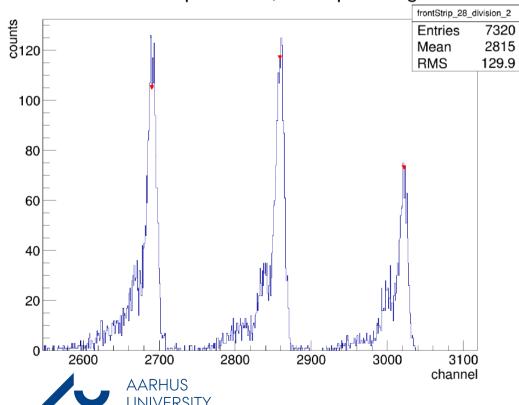


# **Energy calibration**

- Source of <sup>239</sup>Np, <sup>241</sup>Am and <sup>244</sup>Cm
- Either dead layer or internal effect.

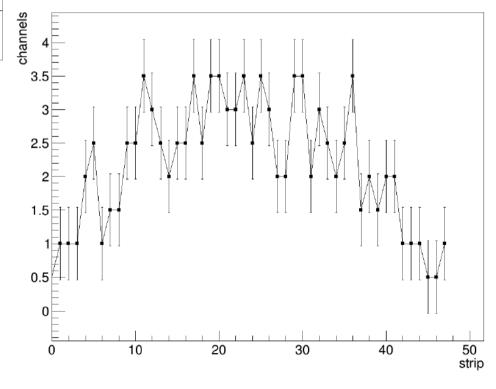






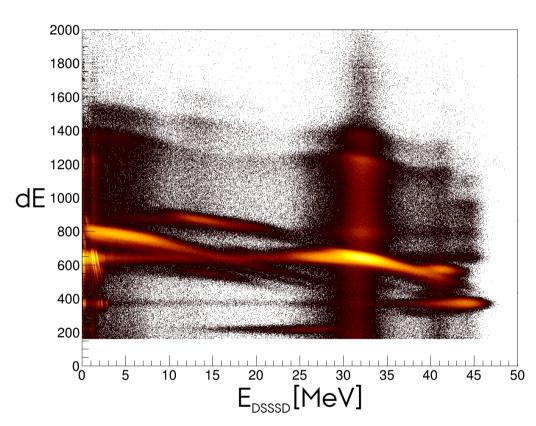
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#### Front strip response curvature, Cm-peak



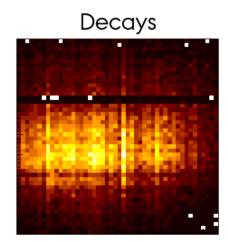
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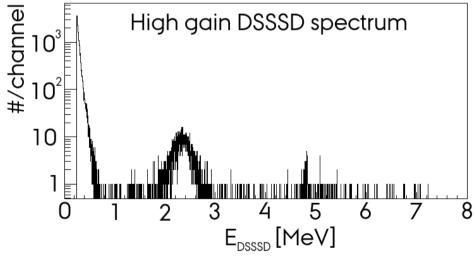
#### The data



•Two amplification chains.







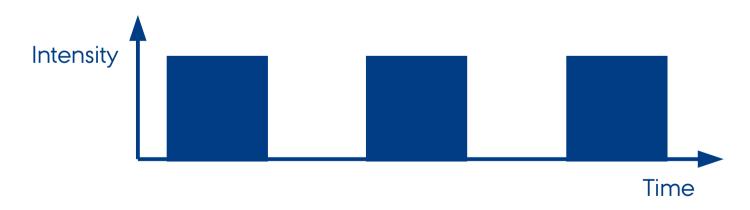


### Normalisation 1, dead time

Trigger → scaler.

$$Eff(\%) = \frac{Accepted triggers}{Total triggers}$$

•Cyclotron duty-cycle.



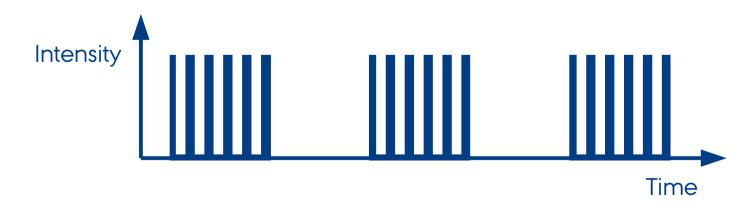


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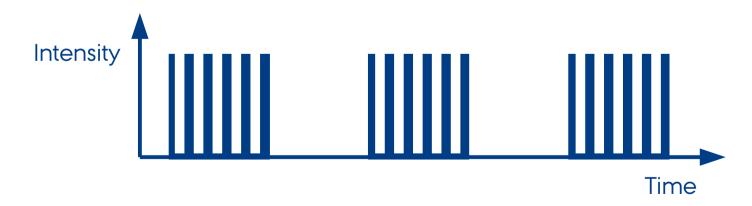


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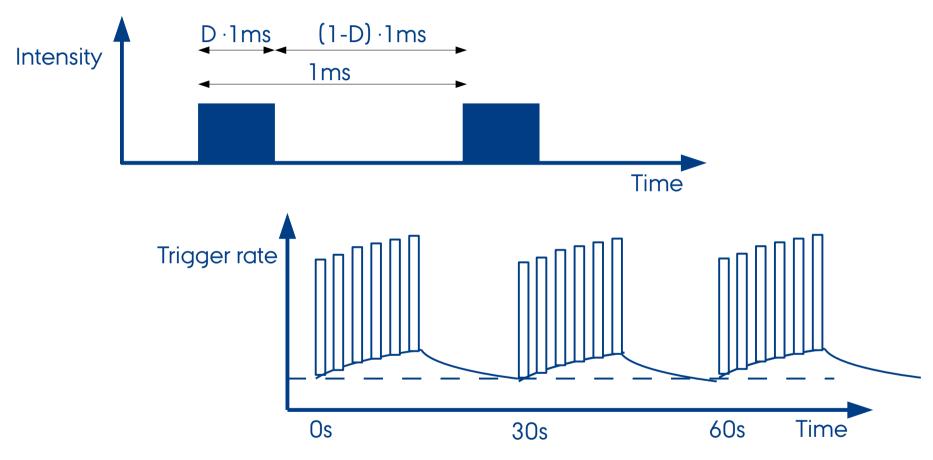
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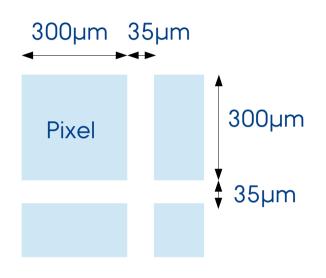
#### Normalisation 1, cont.



Correction now dependent on D (which changes)



## Normalisation 2, sharing



- ~20% of detector area between strips.
  - Implantation vs. decay.
  - To be investigated.

## Summary

- Normalisation is the most challenging part of the analysis.
- Lessons to be learned:
  - Cyclotrons are pulsed.
  - Running with excessive dead time can be bad.
  - Check electronics with oscilloscope.
- Spectroscopy is not affected by dead time.

